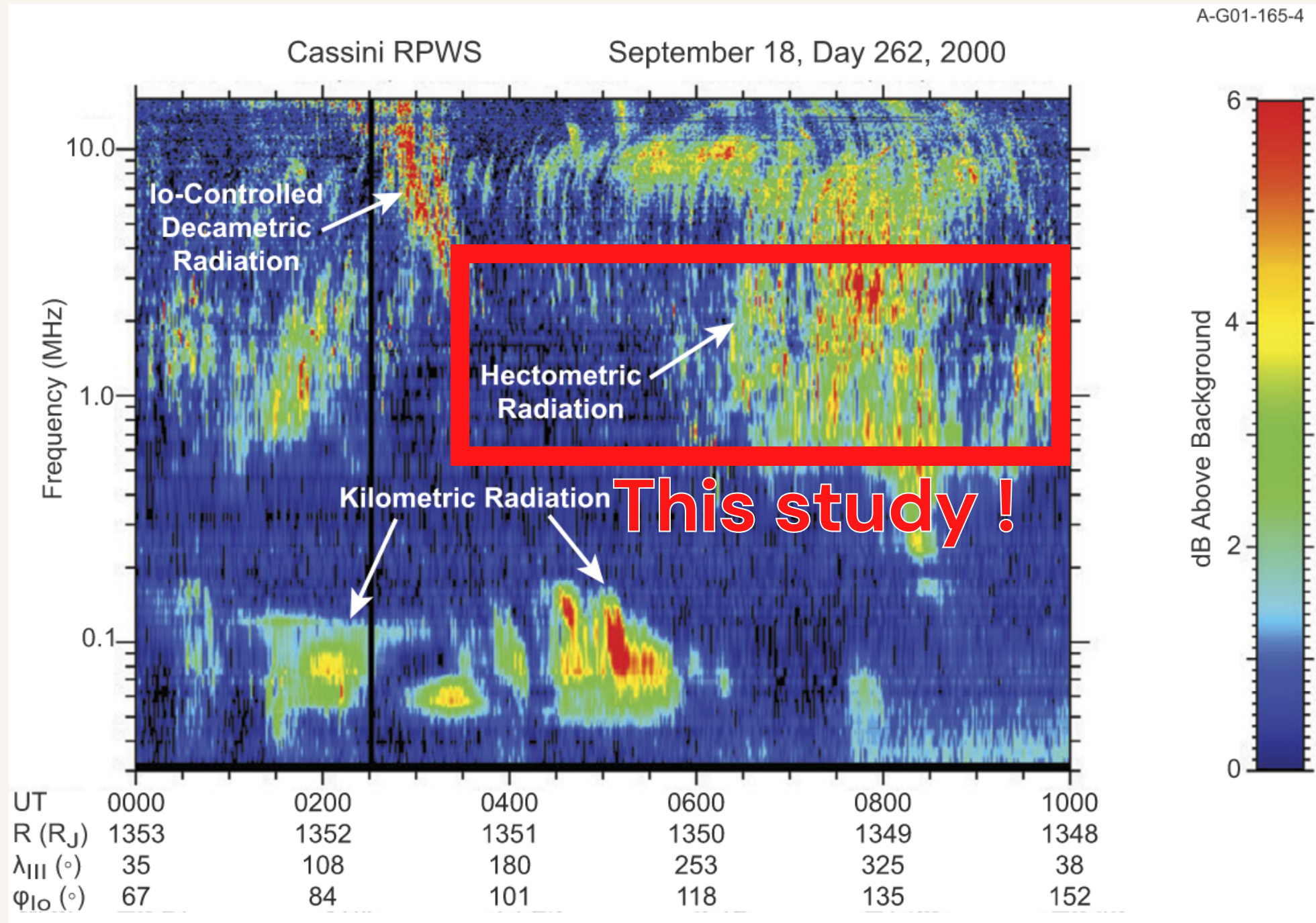


Radio auroral emissions generated by electron beams

B. Collet, L. Lamy, C.K. Louis, P. Zarka, R. Prangé, P. Louarn, W.S. Kurth, F. Allegrini



Credit: Kurth+ (2017)

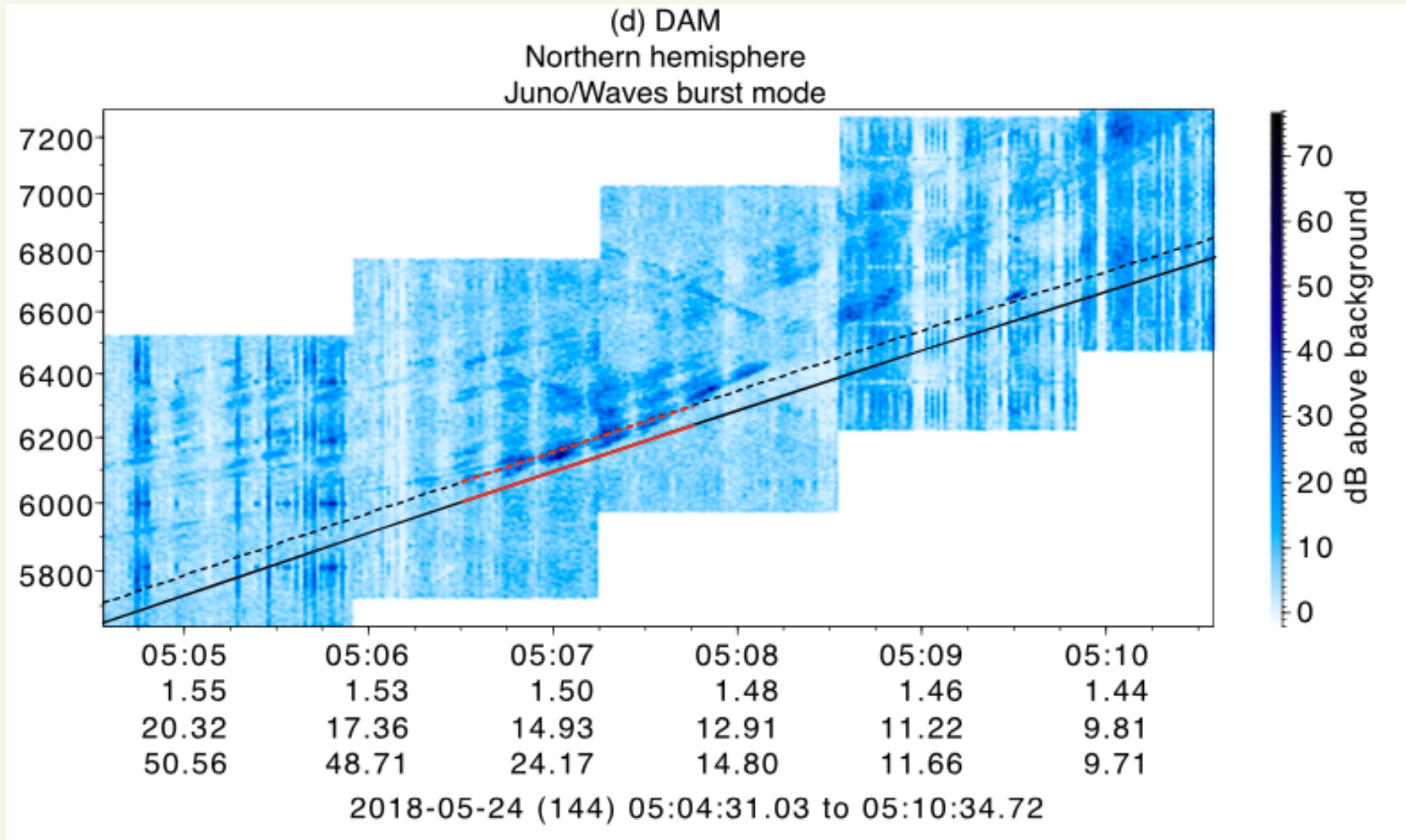
$$\omega = \frac{\omega_{ce}}{\Gamma} + k_{\parallel} v_{\parallel}$$

$$\frac{f_{pe}}{f_{ce}} \ll 1$$

$$\frac{\partial f}{\partial v_{\perp}} > 0 \text{ Population inversion}$$

Radio auroral emissions generated by electron beams

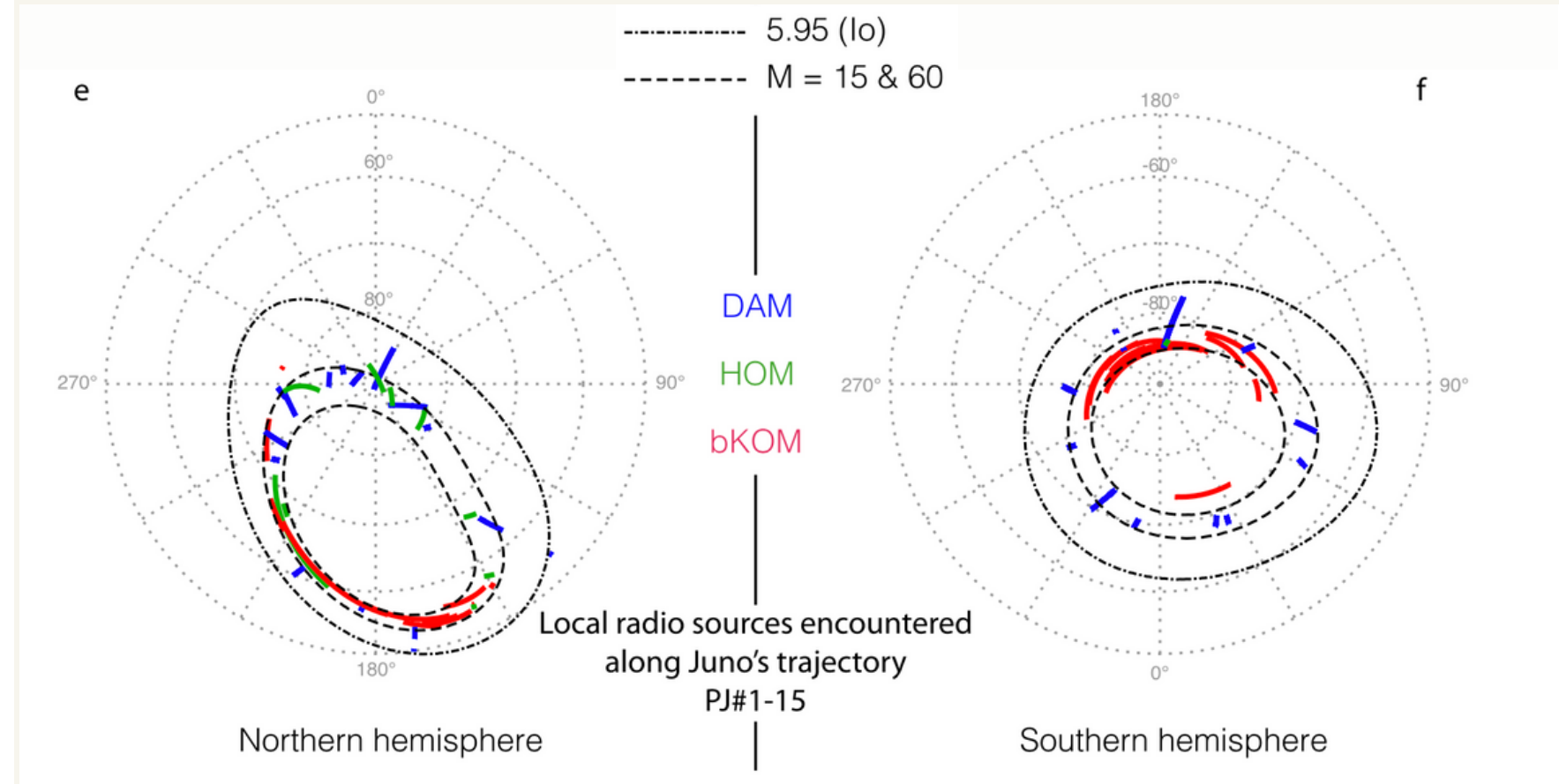
Radio sources



Louis+(2019)

Juno/Waves

Source crossings
identified with $f < f_{ce} + 1\%$



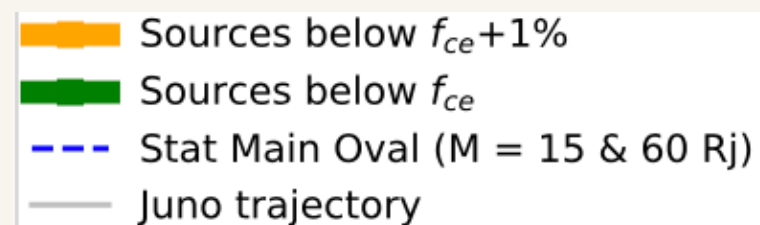
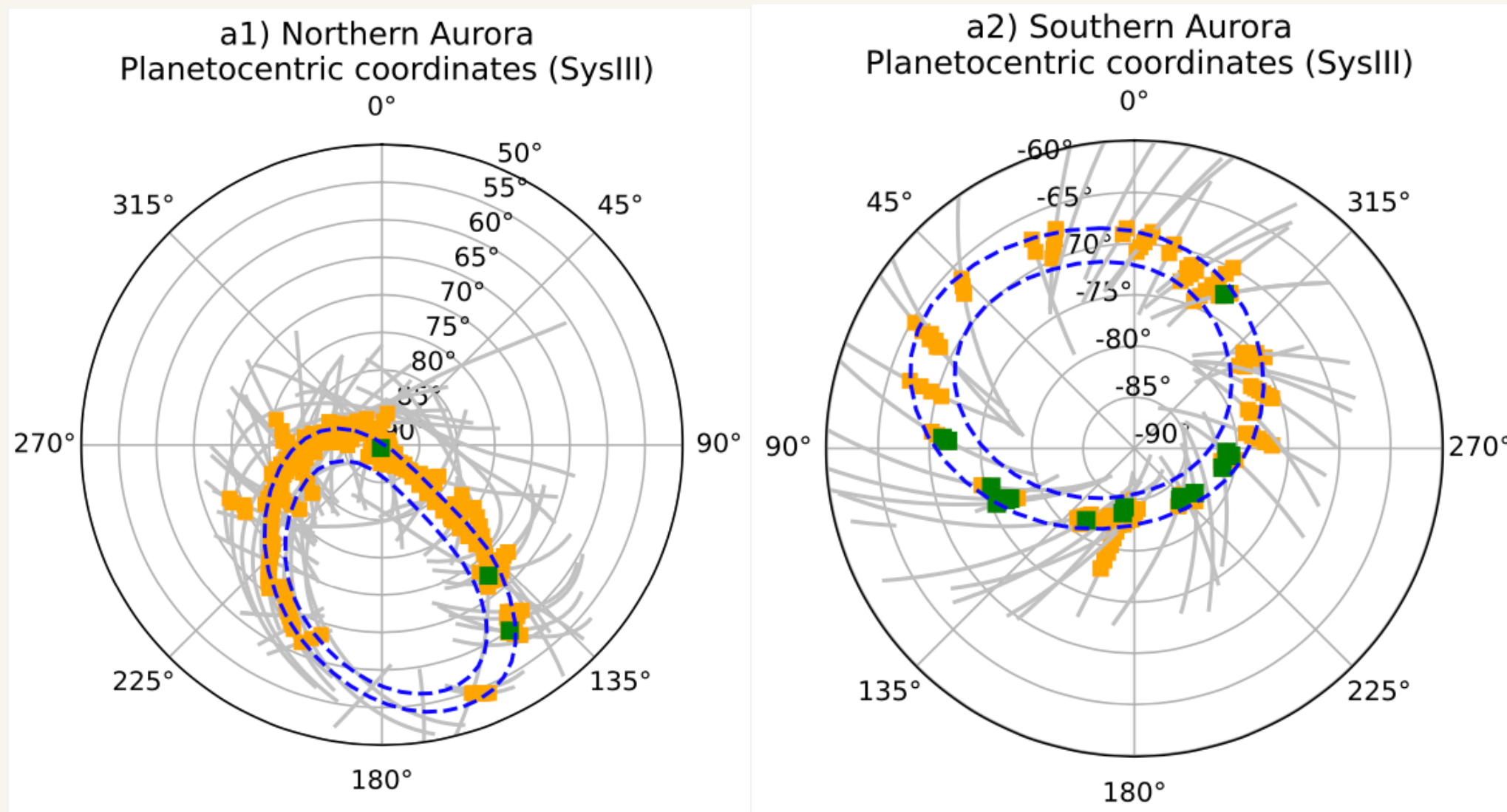
Louis+(2019)

15 first perijoves

$f < f_{ce} ?$

Radio auroral emissions generated by electron beams

Radio survey



Out of the 45 Juno perijoves

- 15 HOM $f < f_{ce}$ sources
- ~90 HOM/DAM $f_{ce} < f < f_{ce}+1\%$ sources

Associated with:

- Dawn side
- Hot plasma
- $f_{pe}/f_{ce} \sim 10^{-3}$
- Depletion of electrons (but no cavity)
- Upward FAC current (Bphi)
- Not systematically connected with brightest UV

Generation ?

Radio auroral emissions generated by electron beams

Cyclotron Maser Instability

$$\omega = \frac{\omega_{ce}}{\Gamma} + k_{\parallel} v_{\parallel}$$

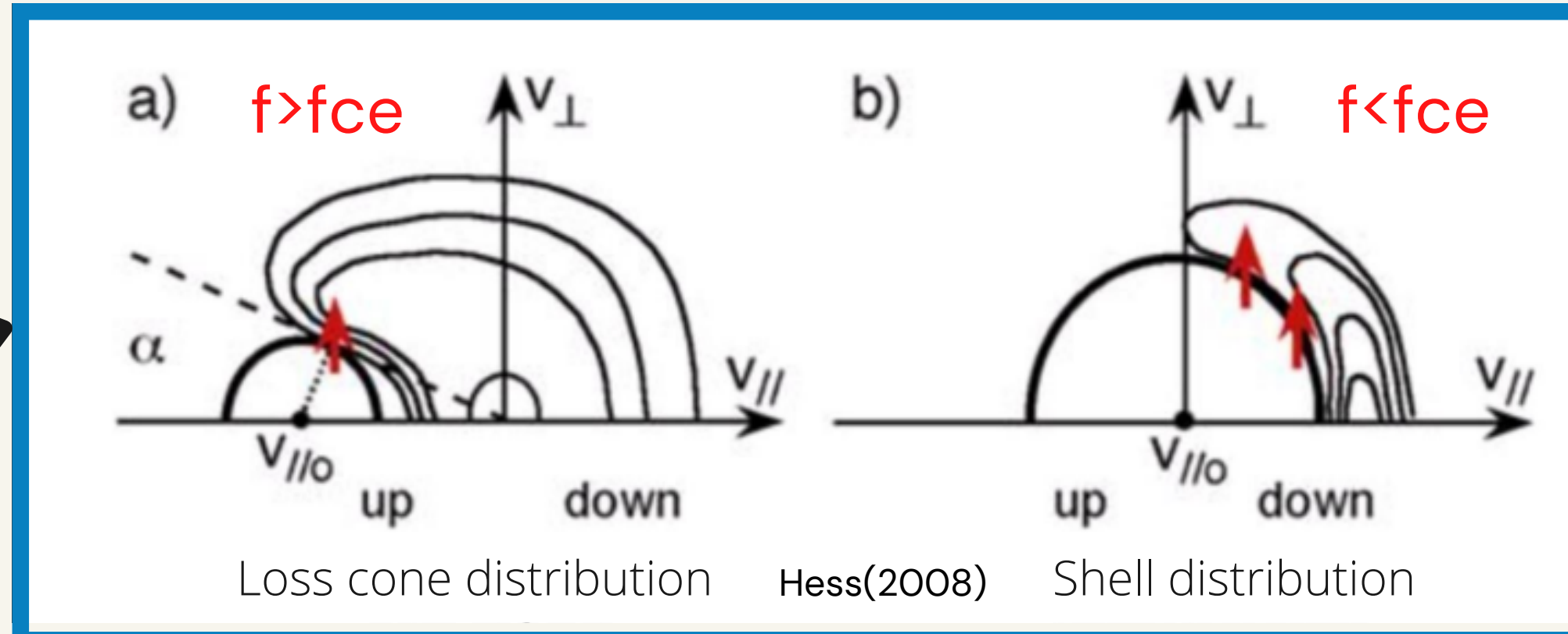
$$\frac{f_{pe}}{f_{ce}} \ll 1 \quad \frac{\partial f}{\partial v_{\perp}} > 0$$

$$v_{\perp}^2 + (v_{\parallel} - v_0)^2 = v_r^2$$

Direction of propagation

Wu&Lee (1979)

Frequency



Identified at Jupiter (Louarn+ 2017)

Main source at Earth

Newly identified at Jupiter

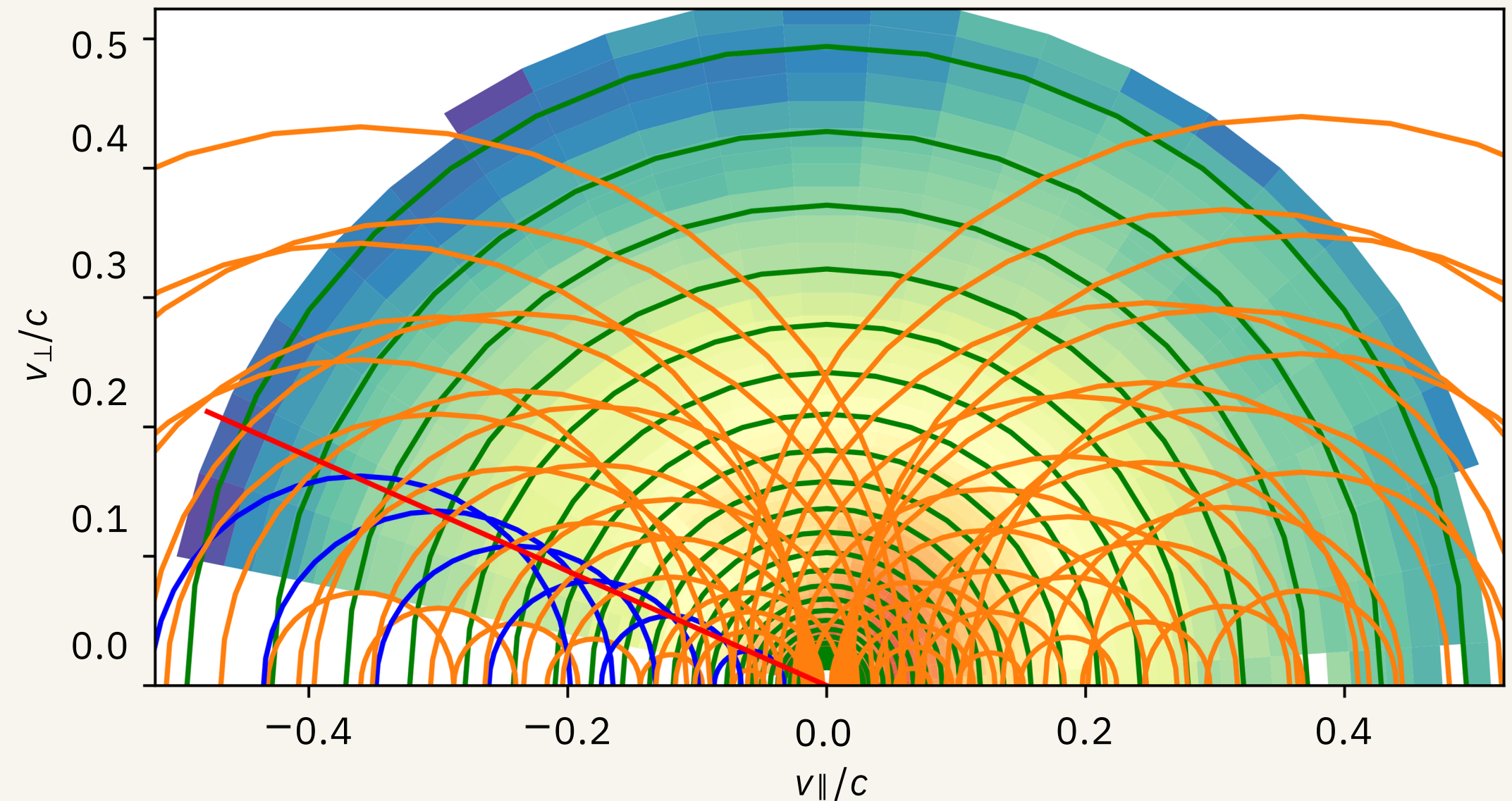
Radio auroral emissions generated by electron beams

Electron measurements

$$\omega_i \propto \oint_{C(v_0, v_r)} \frac{\partial f}{\partial v_{\perp}}$$

3 types of unstable circles :

- **Centered on 0 : shell** $f < f_{ce}$
- **Near the loss cone** $f > f_{ce}$
- **Other (conics, etc...)** $f > f_{ce}$



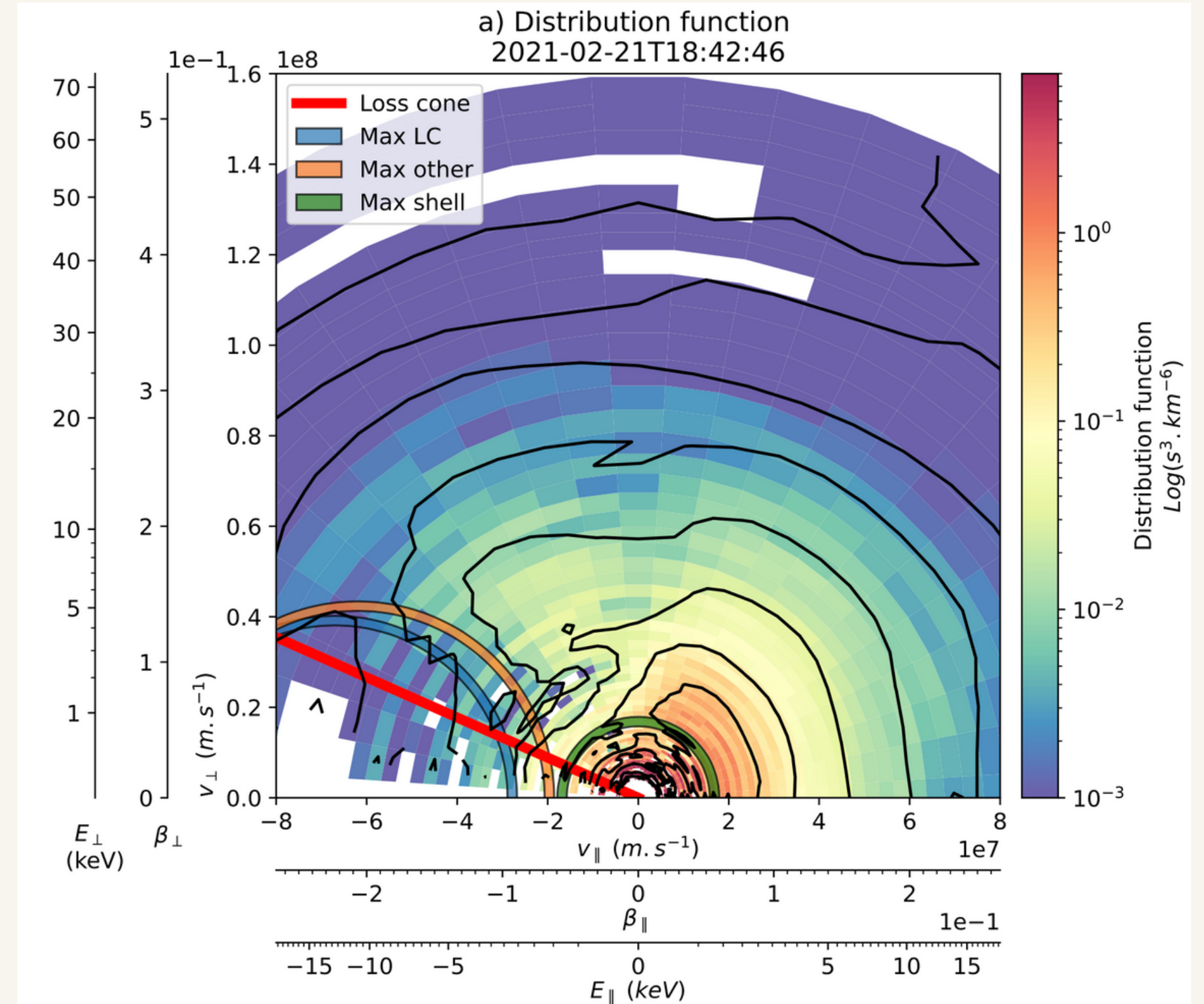
Radio auroral emissions generated by electron beams

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Radio auroral emissions generated by electron beams

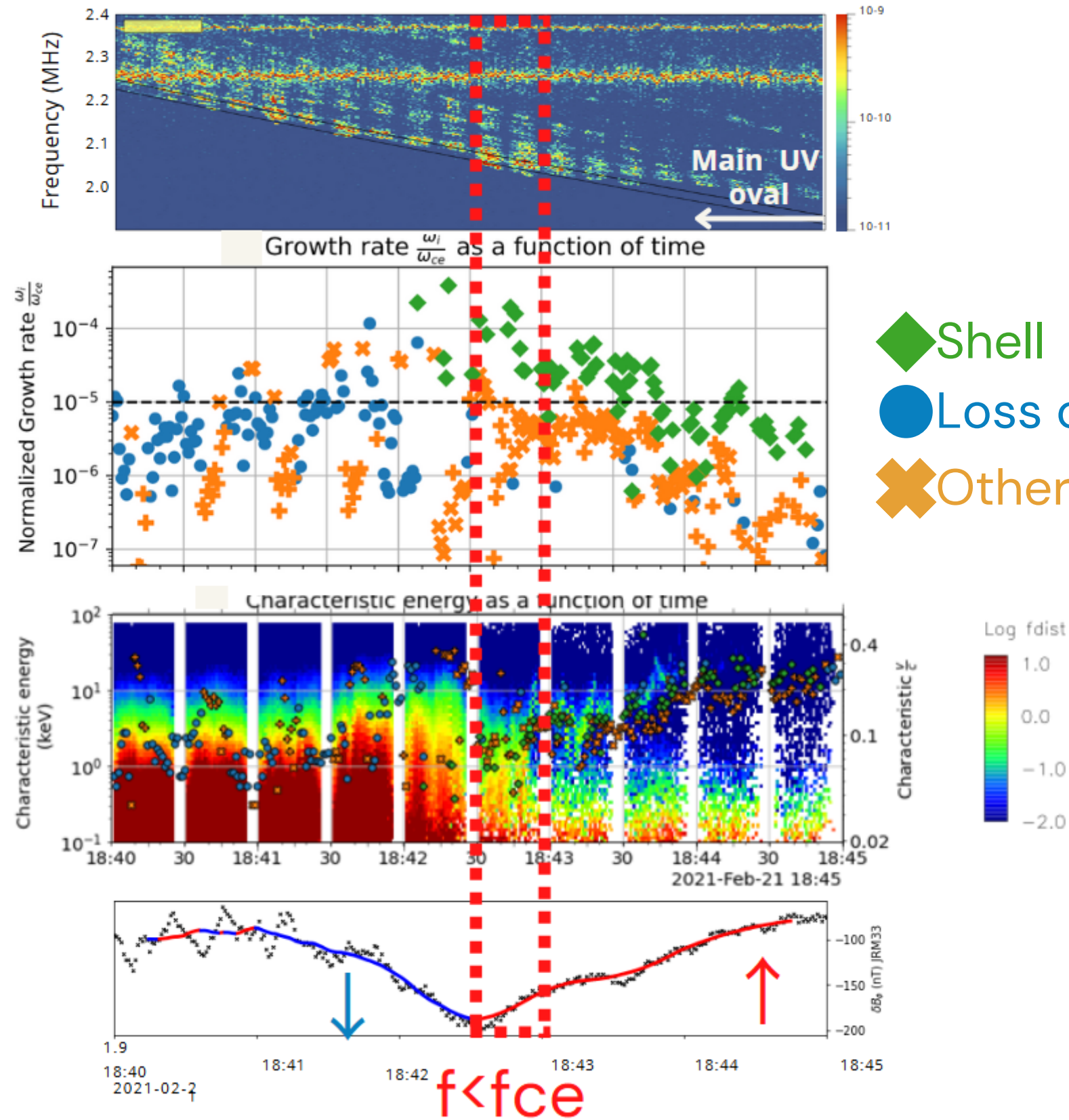
Case study PJ32S

Waves spectra

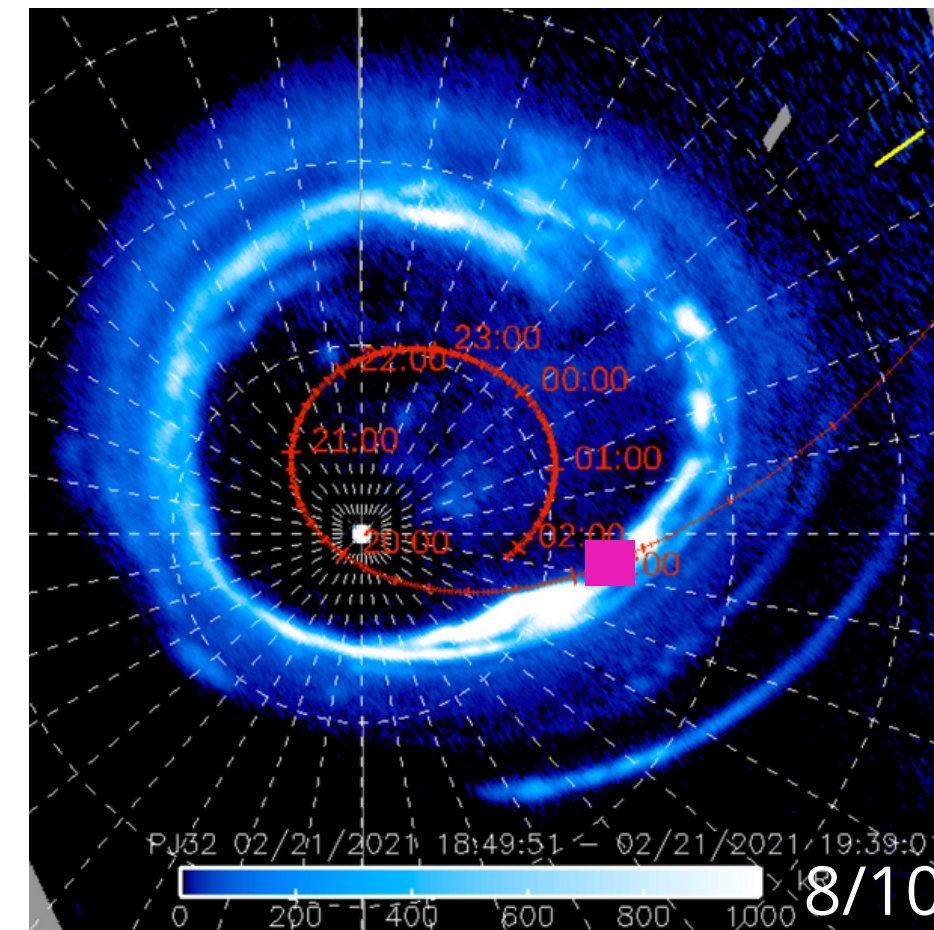
CMI Growth rate

JADE-E \downarrow (1–30 keV)

FAC from Bphi

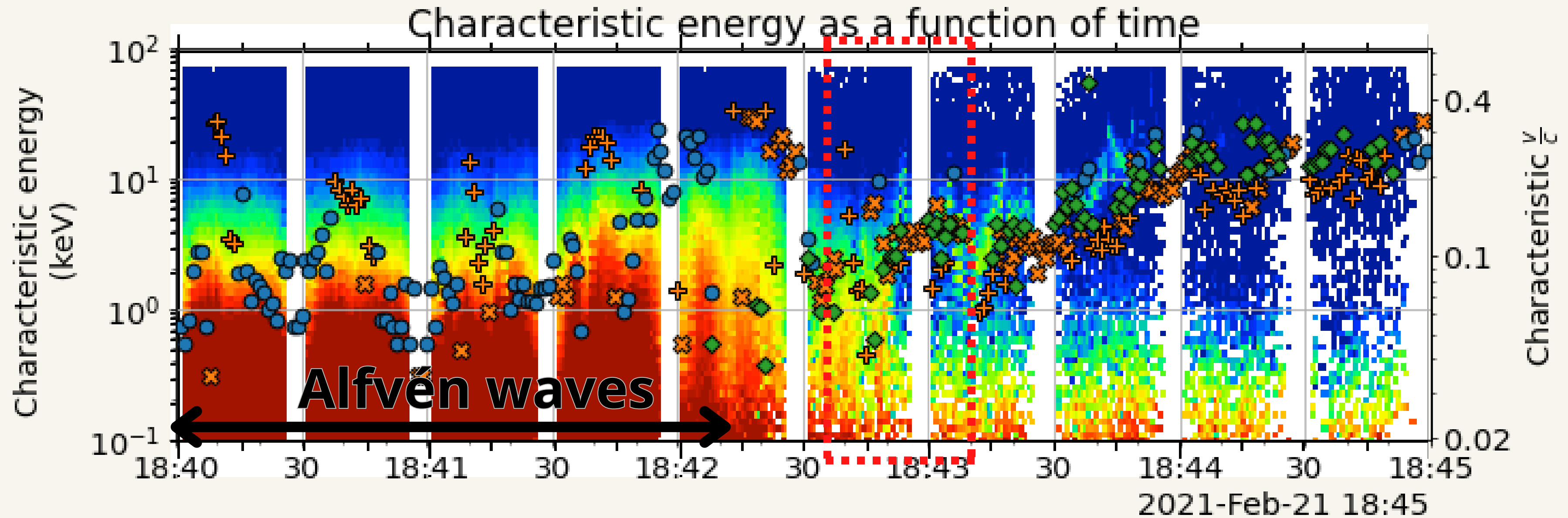


- ◆ Shell
- Loss cone
- ✕ Other (conics or ...)



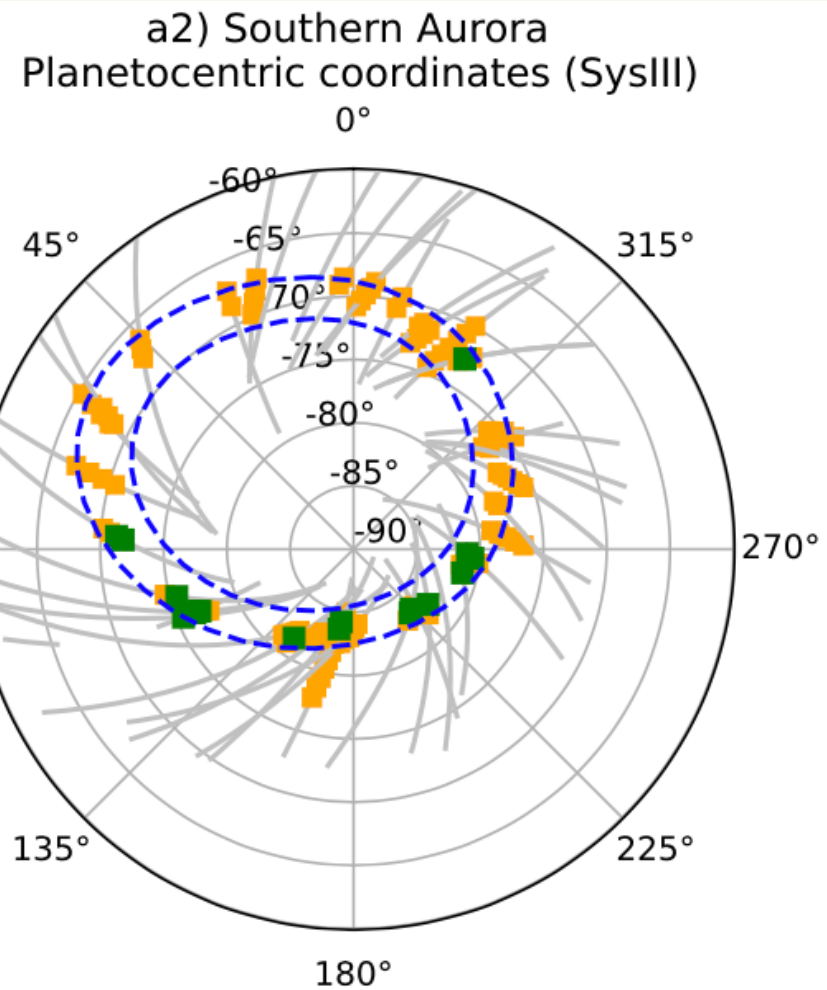
Radio auroral emissions generated by electron beams

Case study PJ32S



Radio auroral emissions generated by electron beams

Results and conclusions



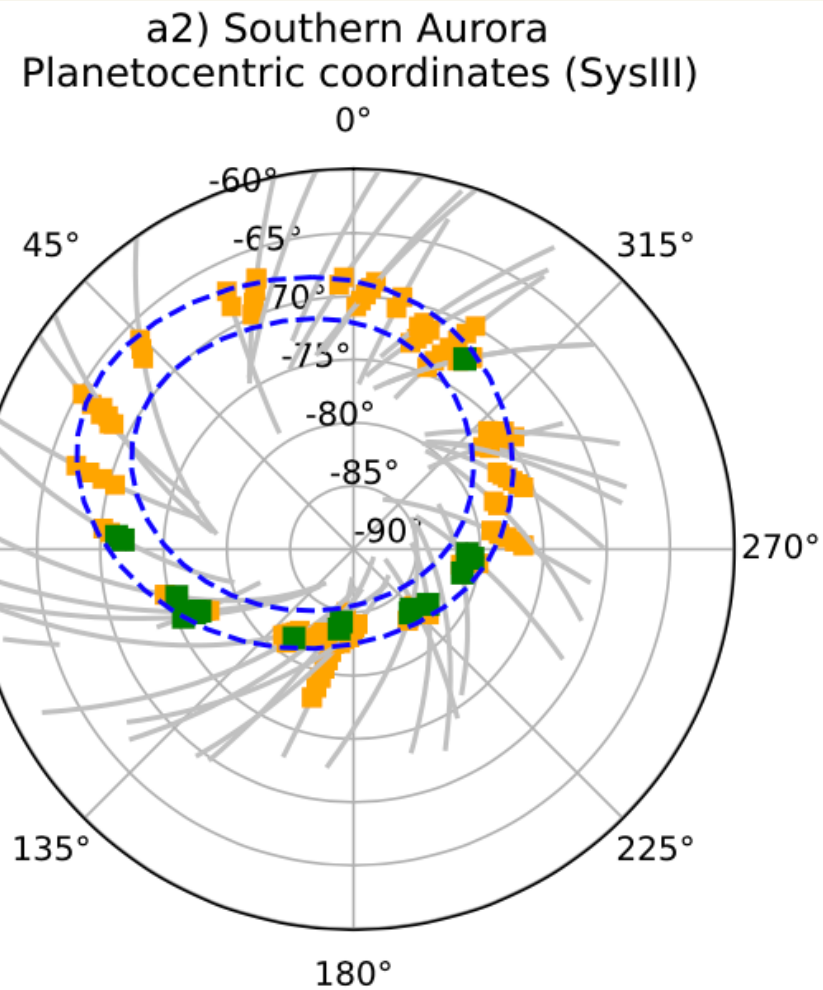
15 HOM f_{ce} candidate sources

=> Associated with **unstable shell EDF**

- Energy lower than at Earth and Saturn (**0.2 - 5 keV**)
- **Mono energetic** signatures (inverted V)
- Associated with enlargement of LC (unstable with same energy !)

Radio auroral emissions generated by electron beams

Results and conclusions



15 HOM $f < f_{ce}$ candidate sources

=> Associated with **unstable shell EDF**

- Energy lower than at Earth and Saturn (**0.2 - 5 keV**)
- **Mono energetic** signatures (inverted V)
- Associated with enlargement of LC (unstable with same energy !)

CONCLUSION

- New source of free energy for HOM in a peculiar plasma environment
- Identification of mono energetic structures with radio

Collet et al.(2024) submitted JGR

Radio survey

15 HOM $f < f_{ce}$ candidate sources

Associated with: Hot plasma ;

- $f_{pe}/f_{ce} \sim 10^{-3}$
- Depletion of electrons (but no cavity)
- Upward FAC current
- Dawn side

What about :-bKOM ?-High energy electrons ?

| PJ | Interval | f (MHz) | $\frac{f-f_{ce}}{f_{ce}}$ | $\frac{f_{pe}}{f_{ce}}$ | UV aurorae | Mono energetic signature | $\frac{n_h}{n}$ | FAC direction | Alfvén waves | Shell energy | Enlargement of LC | Normalized growth rate ($\times 10^{-4}$) | Size (10^3 km) | Estimated intensity ($\text{W}\cdot\text{m}^{-2}\cdot\text{Hz}^{-1}$) |
|-----|------------------------------|---------|---------------------------|-------------------------|------------|--------------------------|-----------------|---------------|--------------|--------------|-------------------|---|-------------------|---|
| 1S | 2016-08-27 13:29:29-30:51 | 5 | -0.6 % | 4×10^{-3} | Poleward. | | 0.5 | ↑ | X | 0.2 | X | 0.3-0.8 | 0.25 | 7.9×10^{-10} |
| 6S | 2017-05-19 06:51:45-52:27 | 3 | -0.2% | 10^{-3} | Diffuse | X | 0.8 | ↑ | | 0.5-3 | X | 0.2 – 2 | 2 | 8.02×10^{-12} |
| 11N | 2018-02-07 12:57:24-36 | 2.5 | -0.3% | 2×10^{-3} | Main Oval | | 0.8 | ↑ | X | 1 | X | 0.5 – 1 | 0.6 | 1.85×10^{-11} |
| 11S | 2018-02-07 14:44:54-45:19 | 2.8 | -0.25% | 3×10^{-3} | Main Oval | | 0.8 | ↑ | X | 2 | X | 0.5 – 3 | 1.25 | 2.56×10^{-11} |
| 21N | 2019-07-21 03:08:05-29 | 2.6 | -0.25% | 2×10^{-3} | Diffuse | X | 0.7 | ↓ | | 3 | X | 0.2 – 2 | 1.25 | 1.20×10^{-11} |
| 24S | 2019-12-26 19:04:04-54 | 1 | -0.3% | 3×10^{-3} | Diffuse | X | 0.15 | ↑ | | 1 | X | 0.2 – 0.8 | 2.5 | 2.69×10^{-11} |
| 25S | 2020-02-17 19:11:43-57 | 1.2 | -0.2% | 3×10^{-3} | Poleward | | 0.5 | ↑ | | 0.2-0.7 | X | 2-3 | 0.75 | 4.13×10^{-11} |
| 28S | 2020-07-25 07:14:06-51 | 2.2 | -0.3% | 2×10^{-3} | Main Oval | X | 0.5 | ↑ | | 2 | X | 0.5 | 1.5 | 4.00×10^{-12} |
| 31S | 2020-12-30 23:08:24-09:20 | 1.1 | -0.25% | 2×10^{-3} | Main Oval | X | 0.6 | ↑↓ | | 2-3 | X | 0.3 – 3 | 2.75 | 8.27×10^{-12} |
| 32S | 2021-02-21 18:42:13-43 | 2.2 | -0.55% | 10^{-3} | Diffuse | X | 0.7 | ↑ | X | 1.5 | X | 0.1 – 1 | 2 | 1.01×10^{-11} |
| 35N | 2021-07-21 7:27:30-40 | 4.0 | -0.2 % | 2×10^{-3} | Diffuse | | 0.7 | * | * | 0.2 | X | 0.25 | 1-10 | 6×10^{-10} |
| 37S | 2021-10-16 18:22:33-25:39 | 1.6 | -0.4% | 3×10^{-3} | Main Oval | | 0.4 | ↓ | | 3 | X | 2 – 20 | (1+)3 | 1.29×10^{-11} |
| 38S | 2021-11-29 15:34:35-36:34 | 1.1 | -0.4% | 2×10^{-3} | Diffuse | | 0.5 | ↑ | | 4 | X | 1 – 20 | (0.5+)1.5 | 8.27×10^{-12} |
| 41S | 2022-04-09 17:05:40-06:35 | 1.3 | -0.3% | 5×10^{-3} | Diffuse | | 0.6 | ↑ | | 5 | X | 2 – 10 | 2.75 | 3.50×10^{-11} |
| 44S | 2022-08-17 16:17:14-34 | 0.9 | -0.8% | 4×10^{-3} | Diffuse | X | 0.5 | ↑ | | 5 | X | 0.2 – 1 | 1 | 2.69×10^{-10} |

$$\omega = \frac{\omega_{ce}}{\Gamma} + k_{\parallel} v_{\parallel}$$

Semi relativistic
electrons



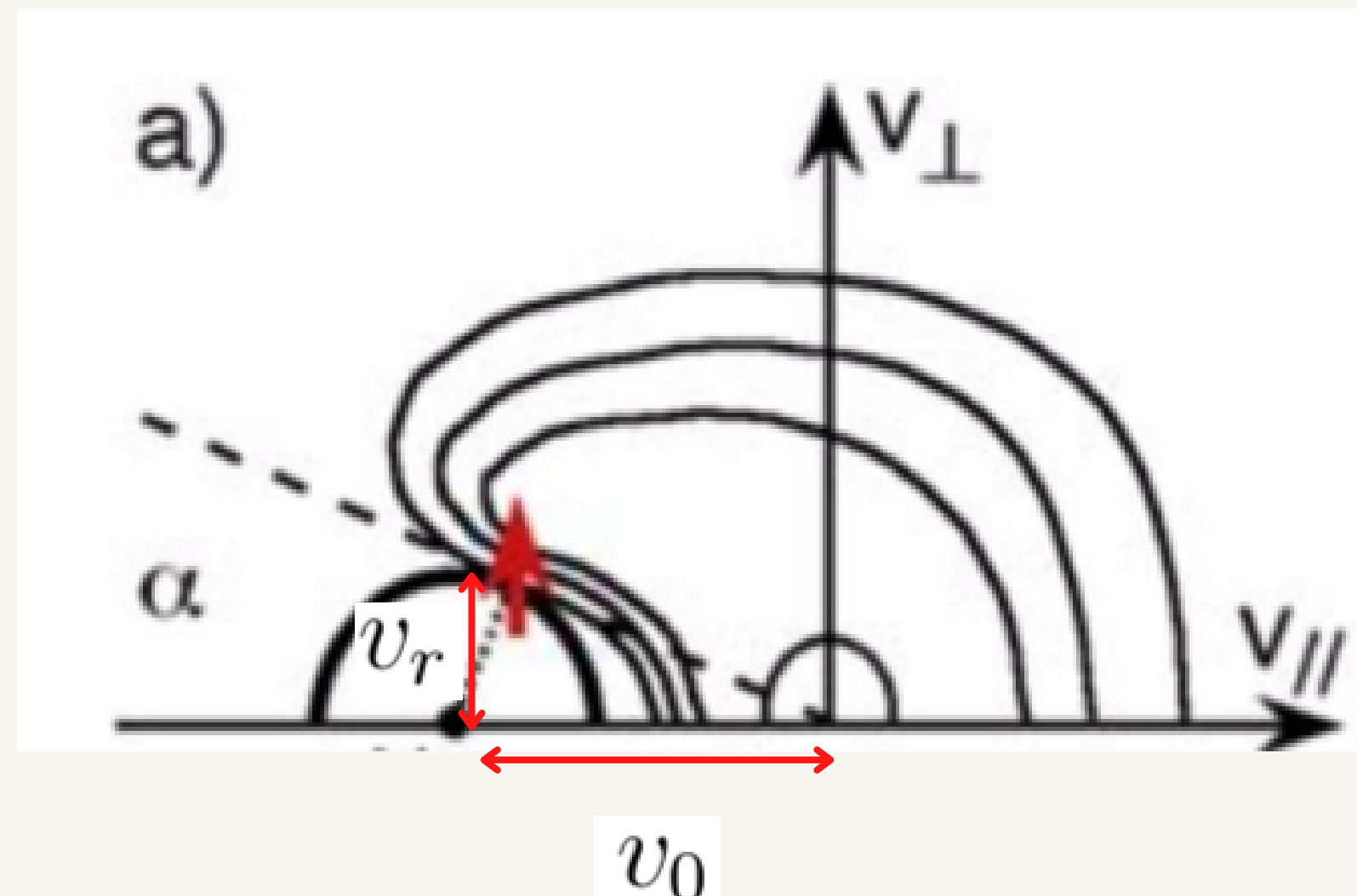
$$v_{\perp}^2 + (v_{\parallel} - v_0)^2 = v_r^2$$

$$v_0 = \frac{k_{\parallel} c^2}{\omega_{ce}}$$

$$v_r = (v_0^2 - 2c^2 \Delta\omega)^{\frac{1}{2}}$$

$$\Delta\omega = \frac{\omega - \omega_{ce}}{\omega_{ce}}$$

- Wave-electron resonance near fce (Wu & Lee 1979)
- Conditions : $\frac{f_{pe}}{f_{ce}} \ll 1$ $\frac{\partial f}{\partial v_{\perp}} > 0$



$$\omega = \frac{\omega_{ce}}{\Gamma} + k_{\parallel} v_{\parallel}$$

Semi relativistic
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- Wave-electron resonance near fce (Wu & Lee 1979)
- Conditions : $\frac{f_{pe}}{f_{ce}} \ll 1$ $\frac{\partial f}{\partial v_{\perp}} > 0$

$$\frac{\omega_i}{\omega_{ce}} = \frac{\left(\frac{\pi\epsilon_h}{2}\right)^2}{1 + \left(\frac{\epsilon_c}{2\Delta\omega}\right)^2} c^2 \int_0^{\pi} d\theta v_r^2 \sin^2(\theta) \frac{\partial f_h}{\partial v_{\perp}}$$

Lien entre
Louarn (2017)
et
Mutel(2007)

Integral
resonance
circle